

***ВЛИЯНИЕ ТОЧНОСТИ СПЕКТРОСКОПИЧЕСКОЙ
ИНФОРМАЦИИ В ЗАДАЧАХ МОНИТОРИНГА
МЕТАНА***

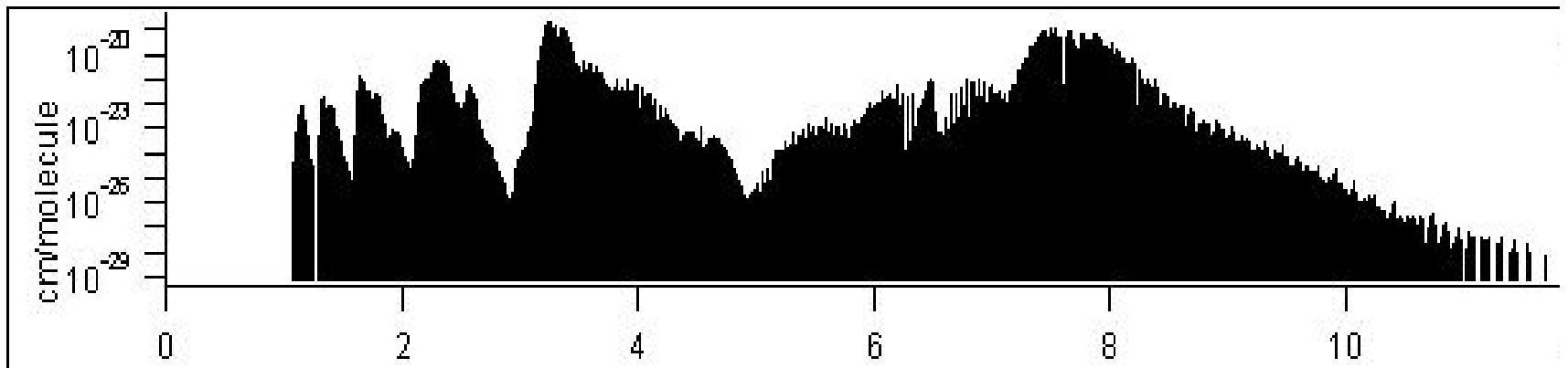
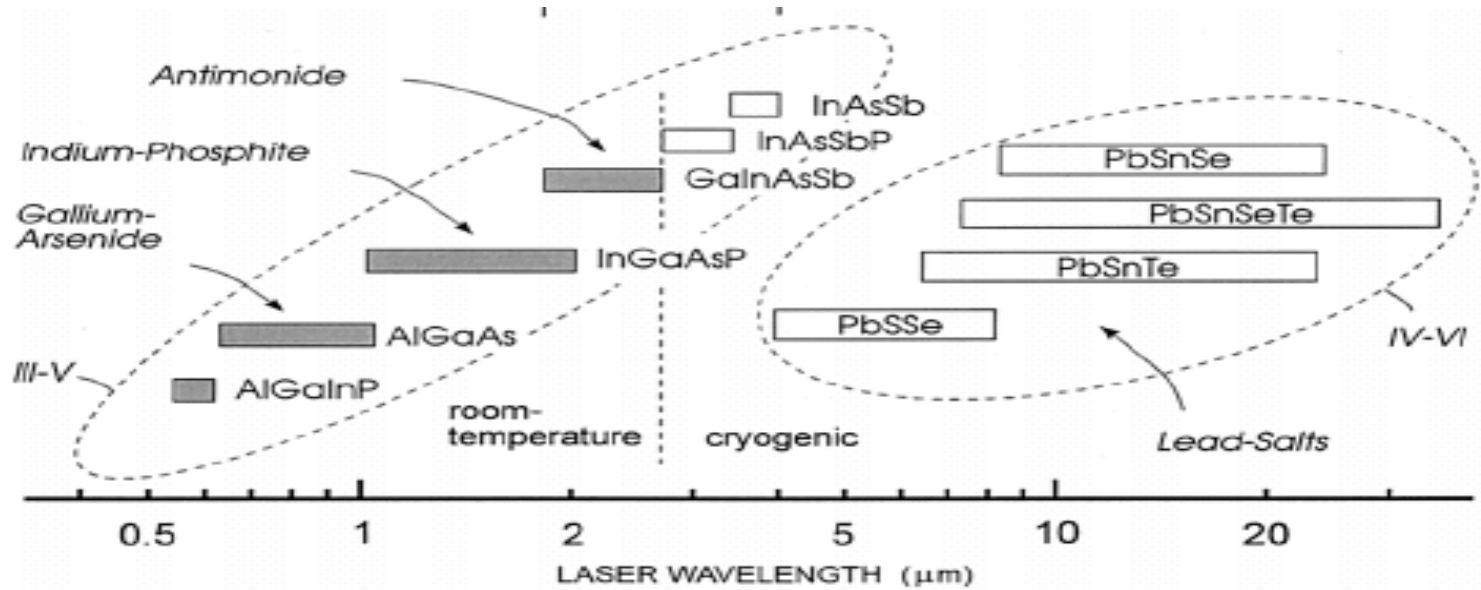
**М.Ю.Катаев, А.В.Никитин
Институт оптики атмосферы СО РАН**


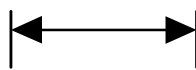
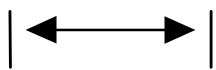
***INFLUENCE OF THE SPECTROSCOPIC
INFORMATION ACCURACY IN TASKS OF THE
METHANE MONITORING***

**M.Yu.Kataev, A.V.Nikitin
Institute of atmospheric optics SB RAS**

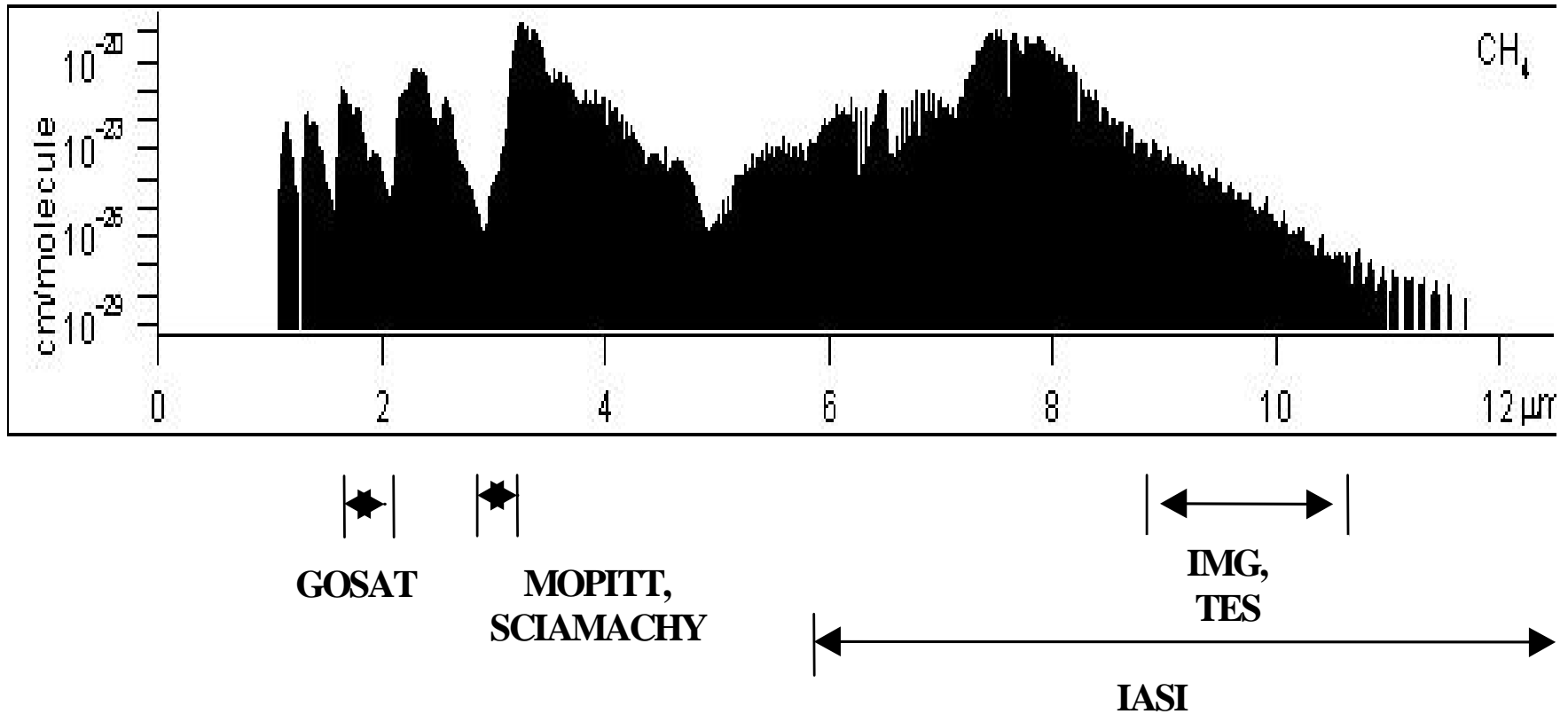
Methane CH₄ is an important trace component in the atmosphere, because of its contribution to global warming as well as its role in complex feedback mechanisms in tropospheric and stratospheric chemistry. Spectroscopic knowledge of the methane spectrum is required for numerous remote sensing applications. Astronomical objects with detectable methane abundances typically include planets, moons and comets in our solar system. The remote control of methane the required absolute high accuracies of the line parameters (0.0001 cm⁻¹ for positions, 1-2% for intensities and broadening). But in modern stage of spectroscopy observations can be taken with better signal to noise and high resolution not all spectral bands. Also the theoretical models are difficult to implement and the spectrum is challenging to interpret, the database for methane has evolved as a mixture of theoretical predictions for the longer wavelengths and incomplete empirical results for the currently intractable regions.

ACTIVE SENSORS FOR METHANE DETECTION




HeNe-laser

**OPG,
CO-laser**

CO2-laser

PASSIVE SENSORS FOR METHANE DETECTION



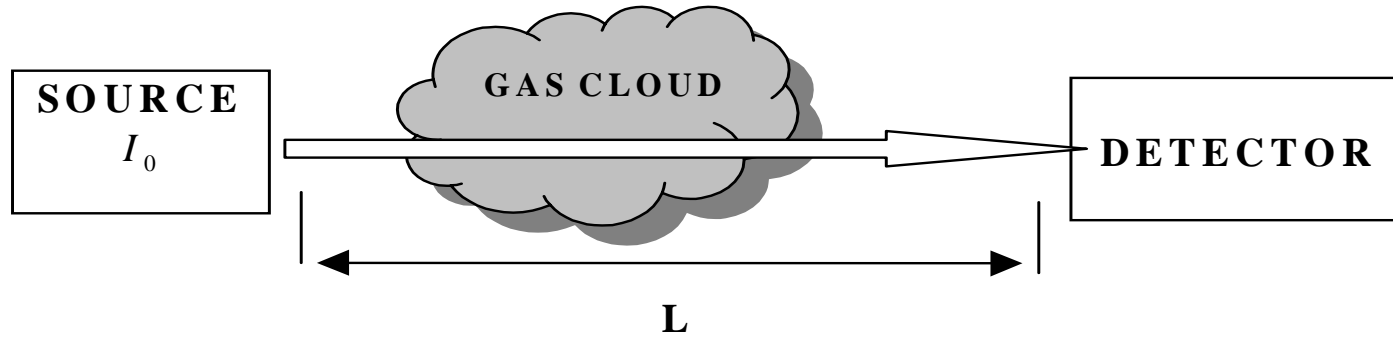
HIGH SPECTRAL RESOLUTION SENSOR: **IMG, TES, SCIAMACHY, GOSAT, IASI**

BROADBAND RESOLUTION SENSOR: **MOPITT**

SPACE-BORNE INSTRUMENTS FOR METHANE DETECTION

Satellite	Device	Method	Profile (P), Total content (TC)	Altitude range km
SHUTTLE	CIRRIS	Limb	P	10-80
UARS	CLAES	Limb	P	10-50
UARS	HALOE	Occultation	P	10-50
EOS-CHEM	HIRDLS	Limb	P	10-50
AQUA	MOPITT	Nadir	TC	-
METOP-1,-2, ESA	IASI	Limb	P	1-30
ADEOS	ILAS	Limb	P	10-60
ADEOS-II	ILAS-2	Limb	P	10-60
ADEOS	IMG	Nadir	TC/P	10-50
UARS	ISAMS	Limb	P	10-80
ENVISAT-1, ESA	MIPAS	Limb	P	5-50
ENVISAT-1, ESA	SCIAMACHY	Occultation, Limb, Nadir	P/TC	5-70
EOS-AM2-3	TES	Nadir, Limb	P/TC	10-60
Spacelab (ATLAS)	ATMOS	Occultation	P	25-60
EUMETSAT	IASI	Nadir	P/TC	1-15

Uniform optical atmospheric path (pressure P , temperature $T \sim \text{const}$)



$$I(\nu) = I_o(\nu) \exp \left\{ -L \cdot \sum_{i=1}^{N_g} K_i(\nu) x_i \right\} \quad (1)$$

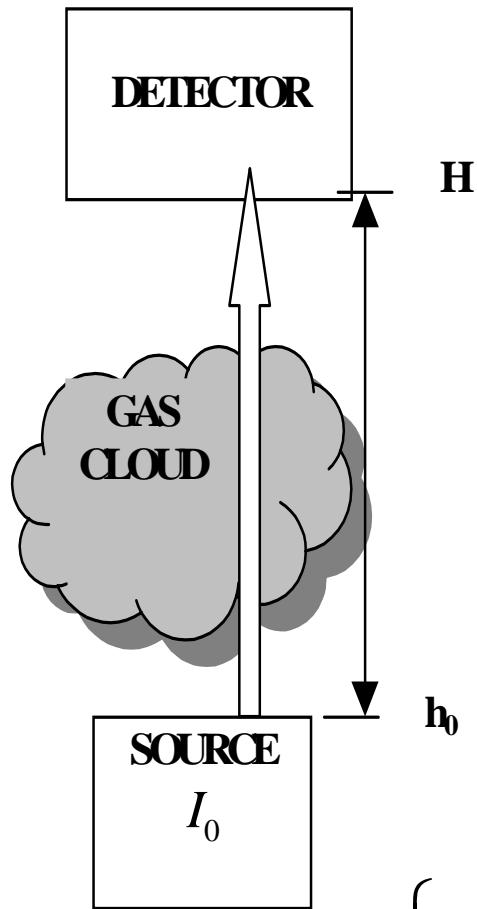
where L is the path length, x is the concentration, N_g is the number of gases in air and K is the absorption coefficient.

Absorption coefficients depends on spectroscopic absorption line parameters (S, α, ν_o):

$$K(\nu) = \frac{1}{\pi} \sum_{j=1}^{N_l} S_j(T) \alpha_j(T, P) / \left[(\nu - \nu_{oj})^2 + \alpha_j^2(T, P) \right]$$

here S is the intensity of absorption line, α is the halfwidth and ν_o is the center of line.

Not uniform optical atmospheric path ($P, T \neq \text{const}$)



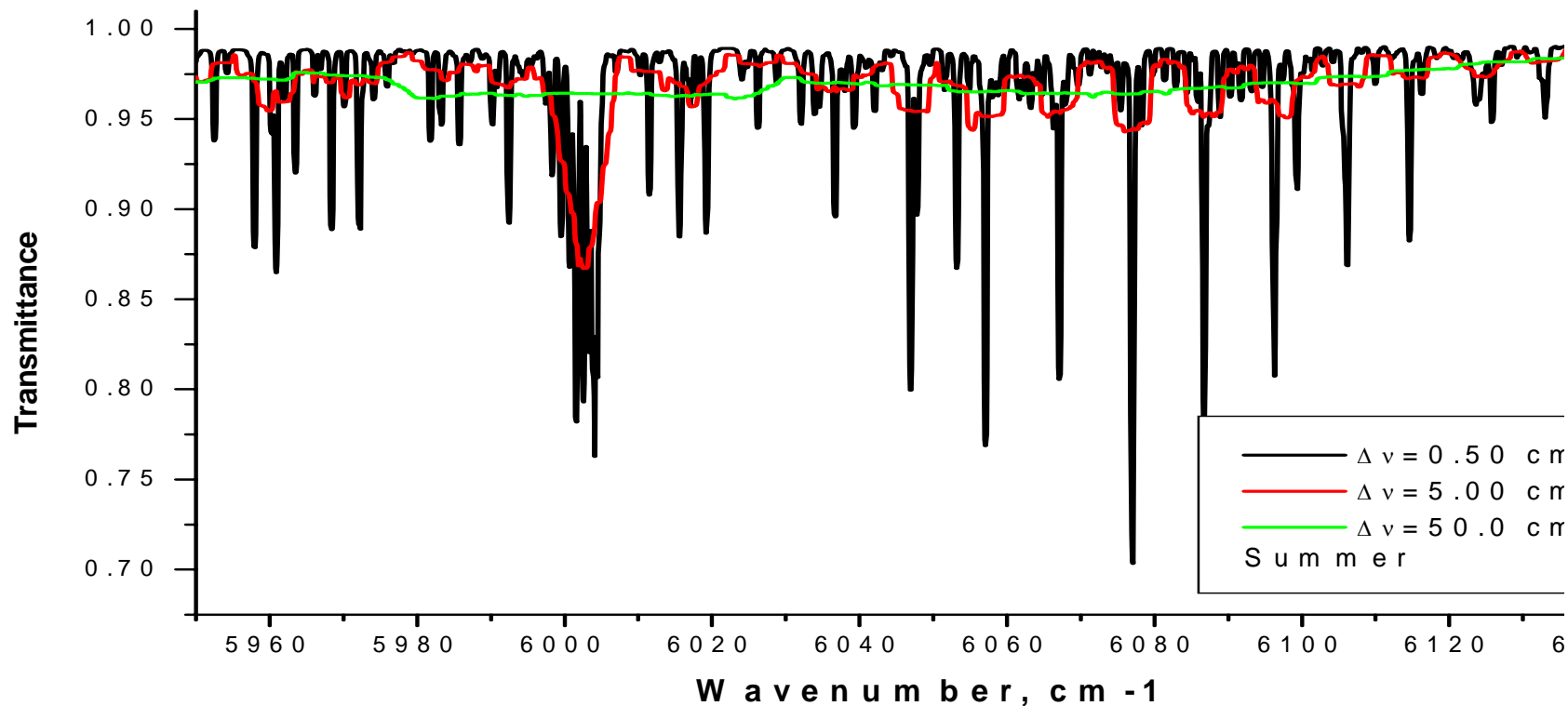
$$I(\nu) = I_0(\nu) \exp \left\{ - \int_{h_0}^H \left(\sum_{i=1}^{N_g} K_i(\nu, h(P, T)) x_i \right) dh \right\} = I_0(\nu) t(P, T)$$

where H, h_0 are the low and high altitude and t is the transmittance.

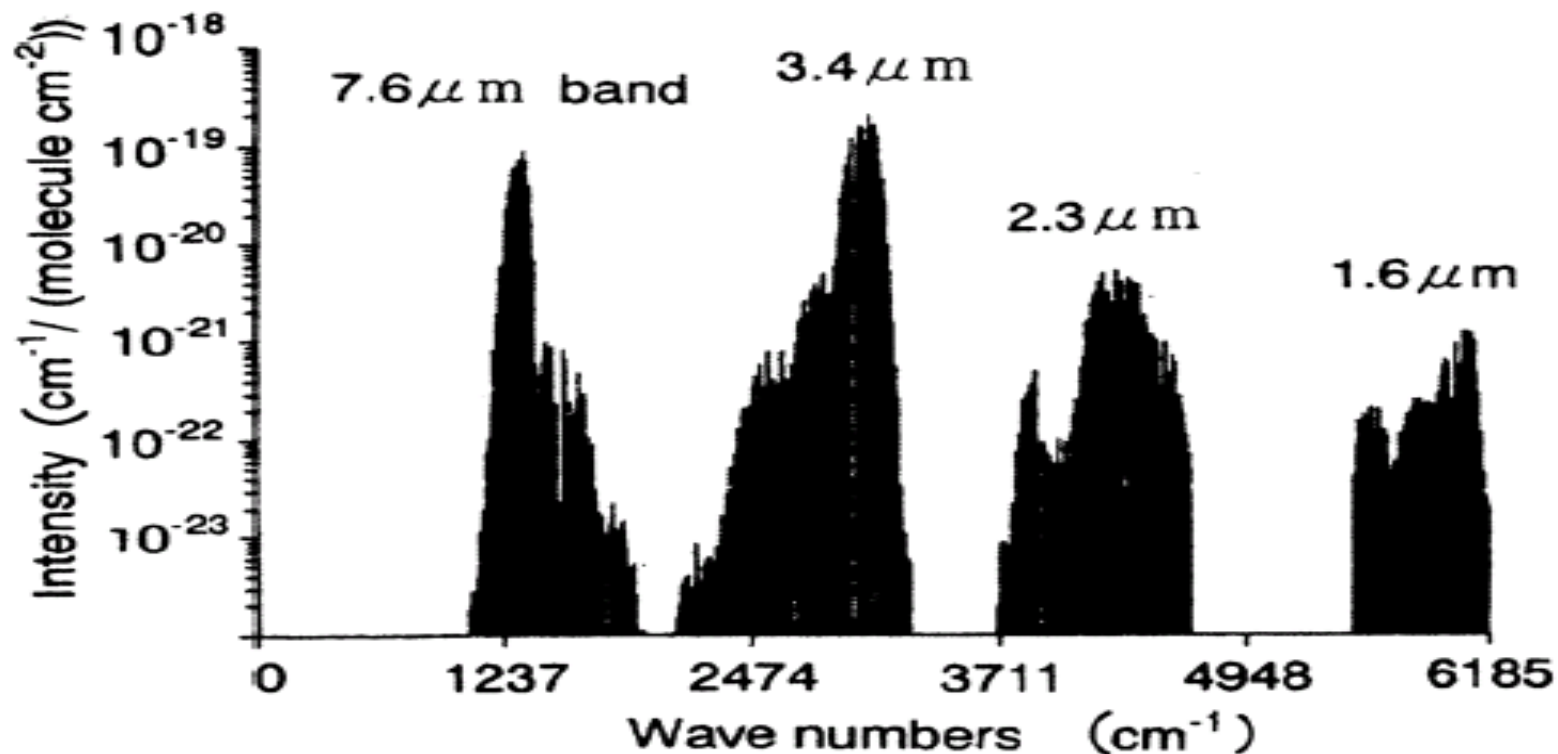
Measured spectral radiance (transmittance) it is connected with true radiance and slit function $A(\nu)$ of the device by a ratio:

$$\hat{I}(\nu) = \int_{\Delta\nu} A(\nu - \nu') I(\nu') d\nu' \quad (3)$$

Equations (1-2) represent a monochromatic kind of radiation, and (3) represent the convolution of the true radiation with the slit function of device.



INFORMATION ABOUT METHANE SPECTROSCOPIC DATA S DATA BASE *HITRAN 2004*



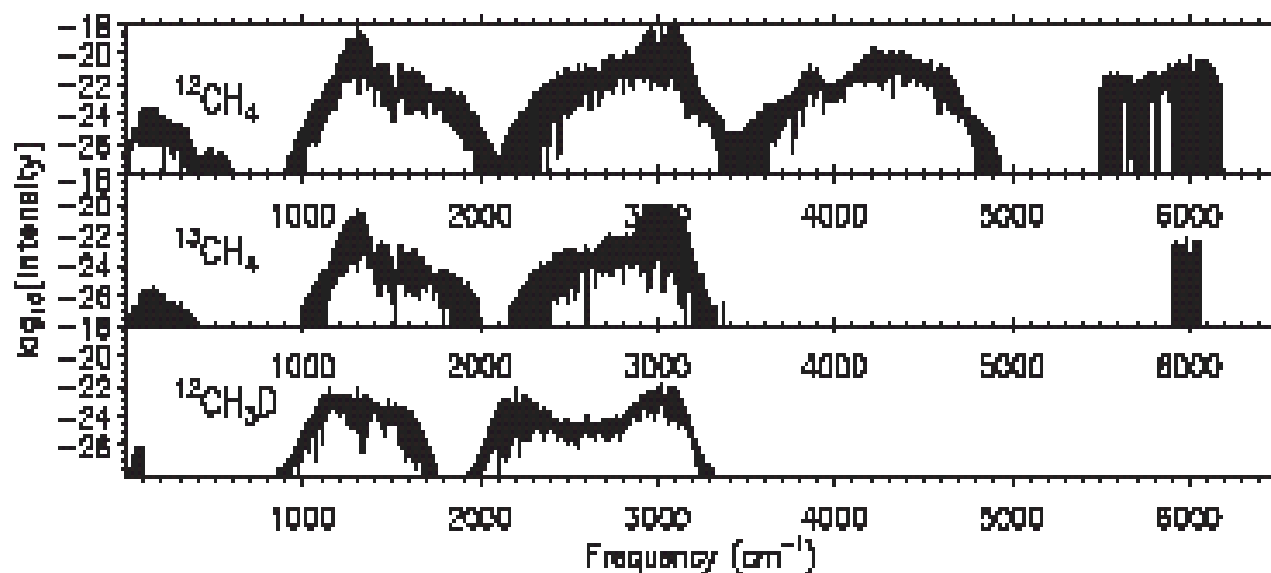
Isotopic abundance used for HITRAN

Molecule	Isotope	Abundance
CH ₄	¹² CH ₄	0.98827
	¹³ CH ₄	0.01110
	¹² CH ₃ D	0.00061

Comparison of 2001 and prior HITRAN methane parameters^a

Polyad	# of Isotopes	# of Bands	Range (cm ⁻¹)	2001 Update		1992–2000 HITRAN	
				Σ Intensity	# Lines	Σ Intensity	# Lines
Rotational	3	8	0–578	5.11×10^{-23}	8681	5.11×10^{-23}	8681
Dyad	3	27	855–2078	5.25×10^{-18}	65,478	5.30×10^{-18}	21,906
Pentad	3	34	1929–3476	1.14×10^{-17}	77,345	1.14×10^{-17}	10,184
Octad	1	9	3370–4810	9.09×10^{-19}	57,332	8.59×10^{-19}	4632
Tetradecad	2	4	4800–6185	1.22×10^{-19}	2632	1.22×10^{-19}	2632

^a Σ Intensity values are in units of cm⁻¹/(molecule cm⁻²) at 296 K.



SPECTROSCOPIC ANALYSIS OF THE MAIN ISOTOPE $^{12}\text{CH}_4$ ABSORPTION BAND

DYAD	PENTAD	OKTAD	TETRADECAD
1200-1600 cm^{-1}	2200-3250 cm^{-1}	3500-4800 cm^{-1}	5200-6700 cm^{-1}
10-4	2*10-3 (4*10-4)	4*10-2 (5*10-3)	-
3%	3% (2%)	15% (5%)	-
			8*10+3 / 200

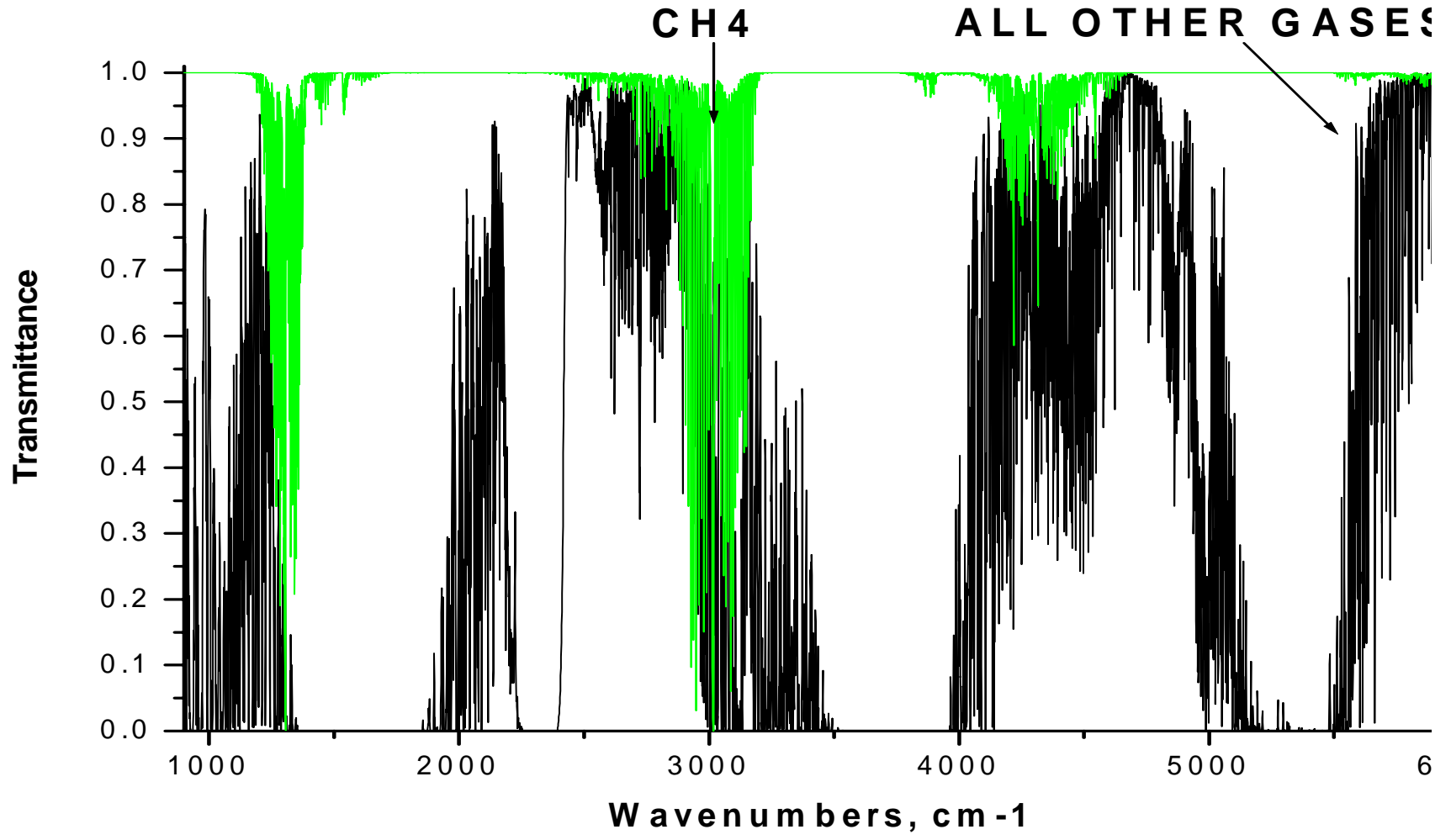
For isotope methane $^{13}\text{CH}_4$ on this day present experiment data up to 5000 cm^{-1} and processing only up to PENTAD.

SPECTROSCOPIC ANALYSIS OF THE MAIN ISOTOPE $^{12}\text{CH}_3\text{D}$ ABSORPTION BAND

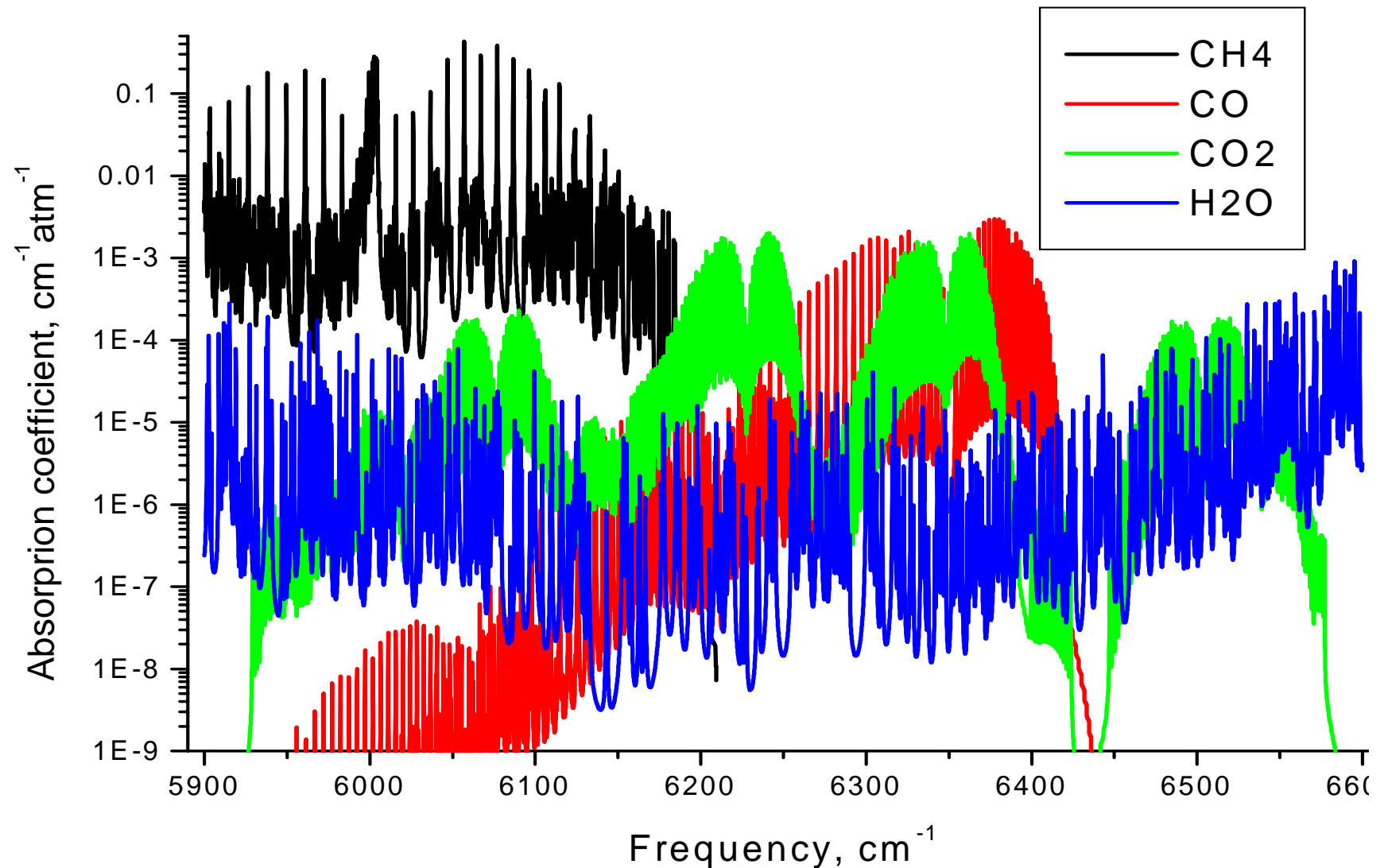
TRIAD	NONAD	HIGH STATES
900-1700 cm^{-1}	1900-3250 cm^{-1}	3250-3700 cm^{-1}
10-3	10-3	- (10-3)
3%	3%	- (4%)

The yellow color is the prospective progress in methane analysis.

CALCULATION TRANSMITTANCE OF THE ATMOSPHERE AND METHANE ON THE OPTICAL PATH 0-100 km



ABSORPTION COEFFICIENTS OF THE GASES ABSORBING SOLAR RADIATION IN 1.6 μm SPECTRAL BAND



ABSORPTION COEFFICIENTS OF THE GASES ABSORBING SOLAR RADIATION IN 10 μm SPECTRAL BAND

